

Caloric curve of 8 GeV/c π^- , \bar{p} + ^{197}Au reactions

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The relationship between nuclear temperature and excitation energy of hot nuclei formed by 8 GeV/c negative pion and antiproton beams incident on ^{197}Au has been investigated with the ISIS 4π detector array at the BNL AGS accelerator. The double-isotope-ratio technique was used to calculate the temperature of the hot system. The two thermometers used, $(p/d-^3\text{He}/^4\text{He})$ and $(d/t-^3\text{He}/^4\text{He})$, are in agreement below $E^*/A \sim 8$ MeV when corrected for secondary decay. Caloric curves derived from successive segments of the H and He kinetic energy spectra show a systematic decrease in temperature as the kinetic energy bin decreases, consistent with “cooling curve” behavior. When extrapolated to the evaporative-peak region, these results provide good agreement with caloric curves measured for similar systems. The caloric curves from this experiment are also compared with the predictions from the SMM multifragmentation model.

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I. INTRODUCTION

Investigation of the thermodynamic properties of hot nuclei formed in nucleus-nucleus collisions is one of the major

objectives of both theoretical and experimental nuclear science. Recently, one of the most interesting and debated questions in this field has been the possibility of a liquid-gas phase transition in finite nuclei [1–5]. In 1995, the ALADIN Collaboration [6,7] presented data suggestive of such a phase transition in peripheral Au+Au collisions at an incident energy of $E/A = 600$ MeV. The caloric curve (the relationship between the temperature and the excitation energy of the hot system), was obtained using the double-isotope-ratio temperature technique [8] with excitation energies reconstructed on an event-by-event basis. The ALADIN results showed a temperature rise with increasing excitation energy per nucleon up to $E^*/A \sim 3$ MeV, followed by a plateau of nearly constant temperature near $T \sim 5$ MeV for excitation energies up to $E^*/A \sim 10$ MeV. For higher excitation energies the initial ALADIN results [6] suggested an increase in temperature, analogous to the heating of the free nuclear gas. These studies were presented as possible evidence for a first-

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order phase transition, encouraging many research groups to investigate the dependence of nuclear temperature on the excitation energy [9–18].

In contrast, the EOS Collaboration data [9–11] obtained for 1 GeV/nucleon Au+C in reverse kinematics, show a monotonic increase of temperature with excitation energy per nucleon with no indication of a plateau. The caloric curves from ISiS experiment E228, 4.8 GeV $^3\text{He} + ^{\text{nat}}\text{Ag}$, ^{197}Au , show an increase in temperature at low E^*/A , a distinct slope change near 2–3 MeV, and a gradual increase in temperature up to $E^*/A = 10$ MeV, but no plateau [12,13]. More investigations of caloric curves have been done [14–19]. However, the experimental conditions varied in all of these measurements, complicating the interpretation of the underlying physics, which will be discussed in more detail below. Whether the liquid-gas phase transition of nuclear matter is a first-order or continuous phase transition remains an important question in understanding the decay of hot nuclei.

Hadron- or light-ion-induced multifragmentation of heavy nuclei provides several unique advantages for the study of the nuclear liquid-gas phase transition. Such collisions can create highly excited nuclei via hard nucleon-nucleon scattering, Δ and other hadronic resonance excitations and pion reabsorption [20–22]. A major experimental advantage of hadron-induced reactions is the formation of a single hot source with a broad population of excited nuclei, permitting studies of the entire caloric curve under identical conditions. Because light ion induced reactions have a wide range of deposited energies, the careful reconstruction of excitation energy on an event by event basis [23] is crucial. In addition, dynamical effects due to collective degrees of freedom, such as rotation, shape distortion, and compression, are minimal in hadron- or light-ion-induced reactions. Therefore, the breakup of the excited system is primarily driven by thermal and Coulomb effects. In this report, the caloric curves for 8 GeV/ $c\pi^-$ and \bar{p} beams incident on a ^{197}Au target have been investigated.

II. EXPERIMENTAL DETAILS AND ANALYSIS

Experiment E900a was conducted at the Brookhaven National Laboratory Alternating Gradient Synchrotron (AGS) accelerator. A secondary negative beam of π^- , K^- , \bar{p} ions of 8.0 GeV/ c momentum was tagged with a time-of-flight and Čerenkov-counter identification system. Beams of about $2\text{--}4 \times 10^6$ particles/cycle, with a cycle time of 4.5 s and a flat top of 2.2 s, were incident on a 2×2 cm² self-supporting 2 mg/cm² thick ^{197}Au target. The beam composition was about 98% π^- , 1% K^- , and 1% \bar{p} at the target. The Indiana Silicon Sphere (ISiS) 4π detector array [24] was used to measure the light-charged particles and intermediate mass fragments from the reaction. The ISiS detector consists of 162 triple-detector telescopes arranged in a spherical geometry. It covers the polar angles from $14^\circ\text{--}86.5^\circ$ in the forward hemisphere and $93.5^\circ\text{--}166^\circ$ in the backward hemisphere with 74% geometrical acceptance. The detector telescopes consist of a gas-ionization chamber, a 500 μm

thick silicon detector, and a 28-mm CsI scintillator with photodiode readout. The gas-ionization chambers provide charge identification for the kinetic energies as low as 1.0 MeV/nucleon. Isotope identification was possible for LCPs (LCPs = H and He) with laboratory kinetic energy per nucleon above $E_k/A \geq 8$ MeV. The analysis involved 2.5×10^6 π^- and 2.5×10^3 \bar{p} events that met the detector array trigger multiplicity requirement ($M_{\text{th}} \geq 3$) for thermal particles. Further experimental details can be found in [25–27].

In order to investigate the heating curve for hot nuclei formed in the 8 GeV/ $c\pi^-$, $\bar{p} + ^{197}\text{Au}$ reactions, the excitation energy (E^*) for the thermal source (defined below) has been calculated on an event-by-event basis according to the prescription

$$E^* = \sum_{i=1}^{M_{\text{th}}} K_i + M_n \langle K_n \rangle - Q + E_\gamma. \quad (1)$$

Here K_i is the kinetic energy for each thermal charged-particle in an event of thermal multiplicity M_{th} . M_n is the thermal-like neutron multiplicity, which was estimated from the neutron-charged particle correlations reported for 1.2 GeV/ $\bar{p} + ^{197}\text{Au}$ reactions measured by Goldenbaum *et al.* [28]. The mean neutron kinetic energy is taken from the correlation between $\langle K_n \rangle$ and E^*/A and then Eq. (1) is iterated to obtain self-consistency. Q is the binding energy difference in the reconstructed event and E_γ is a small term to correct for energy released in photons, which is assumed to be $E_\gamma = M_{(Z \geq 3)} \times 1$ MeV. The procedures for calculating E^*/A are described in detail in Ref. [23], along with discussions of the effects of fluctuations and uncertainties in the procedure. Additional discussion can be found in Refs. [25,29].

The thermal particles used in the calculation of the excitation energy are separated from the fast-cascade/preequilibrium contributions based on their kinetic energy in the average source frame, as determined from a two-component moving-source fit to the data [12,13]. The selection of thermal particles (energy gate for thermal particle selection for E^* calculation) E_{th} makes use of the analysis of data from the 1.8–4.8 GeV $^3\text{He} + ^{197}\text{Au}$ reaction [12,13]

$$E_{\text{c.m.}} \leq 30 \text{ MeV}; \quad Z = 1 \quad (2)$$

and

$$E_{\text{c.m.}} \leq 9Z + 40 \text{ MeV}; \quad Z \geq 2. \quad (3)$$

From this selection, only the thermal-like ejectiles are included in the calculation of the excitation energy.

III. RESULTS

A. Constructing the caloric curve

In Fig. 1 the kinetic energy spectra for hydrogen and helium ions from the 8 GeV/ $c\pi^- + ^{197}\text{Au}$ reactions transformed to the source frame are shown at a forward and a backward angle, and at excitation energies of 4 MeV/nucleon and 8 MeV/nucleon. The very low energy thresholds are achieved by using the gas-ionization chamber/silicon detec-

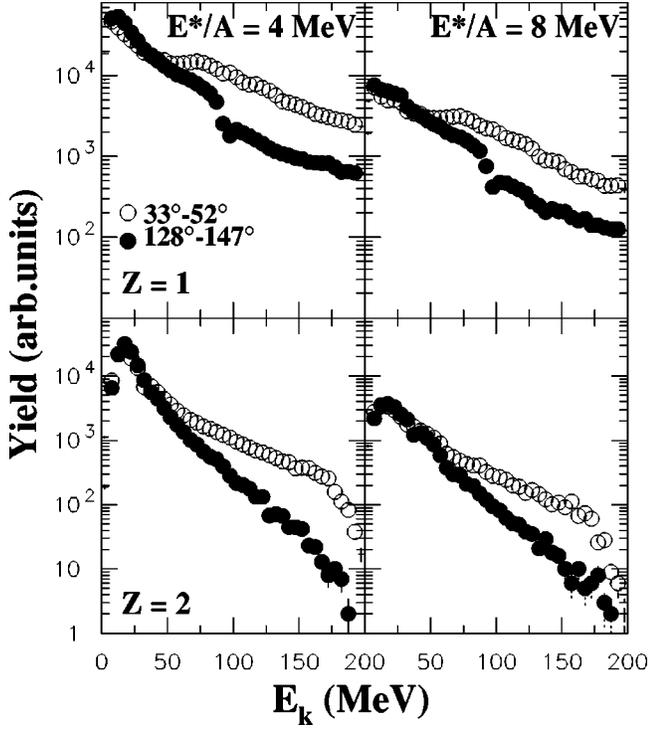


FIG. 1. Forward- and backward-angle kinetic energy spectra in the source frame for $Z=1$ (top panels) and $Z=2$ (bottom panels) at $E^*/A=4$ MeV (left panels) and $E^*/A=8$ MeV (right panels). Statistical uncertainties are shown as error bars but most are smaller than the symbols.

tor stage of the telescopes. The discontinuity near 90 MeV in the H spectra is due to detector punch-through effects. For higher energy $Z=1$ particles (above 90 MeV), energies were determined from the energy loss in the CsI crystal. This permitted measurement of “gray particles” (~ 100 – 1000 MeV unidentified charged particles emitted during the fast cascade stage of the reaction) up to about 350 MeV, assumed to be primarily protons, based on their energy loss in the CsI. The importance of fast-cascade/preequilibrium emission can best be seen in the forward angle $Z=2$ spectra. For kinetic energies above about 60 MeV, the forward angle spectra show a transition to flatter slopes, indicative of preequilibrium emission. In the comparison between the higher and lower excitation energies, the higher excitation energy spectra have flatter slopes in the low energy (thermal) portion of the spectra, as would be expected for a system at a higher temperature. This difference can be seen most clearly in the backward angle $Z=2$ spectra.

ISiS provides isotope resolution for particles with kinetic energies $E_k/A \geq 8$ MeV, primarily $Z=1$ – 3 isotopes in these experiments. In Figs. 2 and 3 the spectra for LCPs (LCP = H and He isotopes) are shown for both a forward (33° – 52°) and a backward (128° – 147°) angle range at two excitation energies for the emitting source, $E^*/A=4$ and 8 MeV. In all cases the slopes of the spectra in the region of interest, $E_k \sim 30$ – 50 MeV (see below), become systematically steeper with increasing LCP mass. This behavior is consistent with the EES model [30], which predicts that the fragment emission time increases with the mass of the fragment; i.e., light

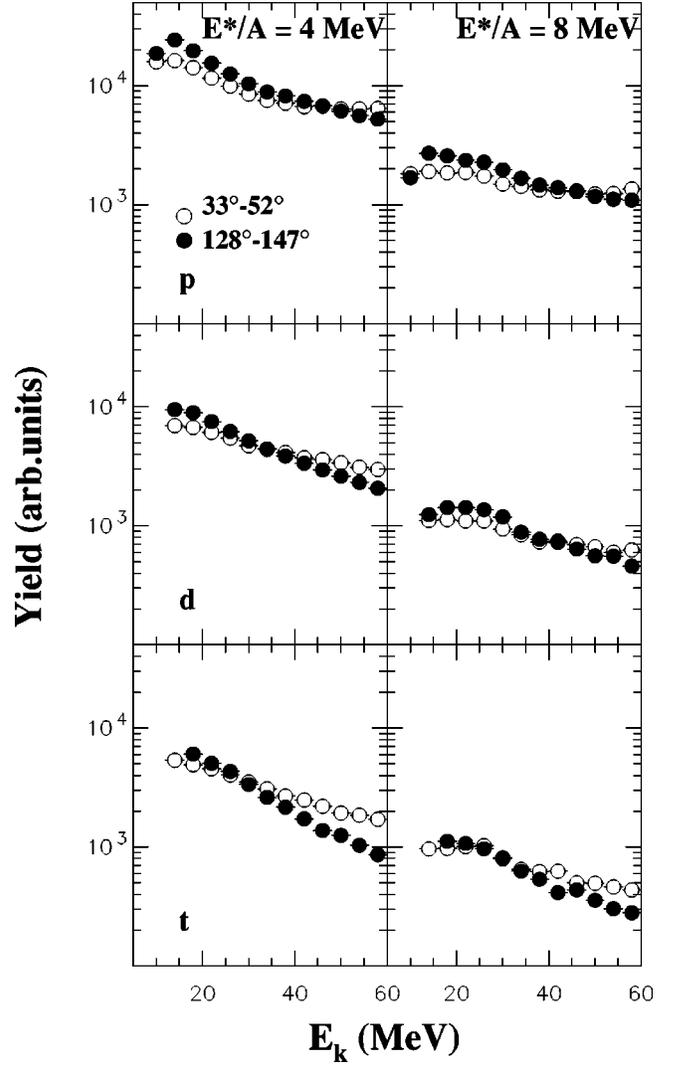


FIG. 2. Forward- and backward-angle kinetic energy spectra for p , d , and t at $E^*/A=4$ MeV (left panels) and $E^*/A=8$ MeV (right panels). Statistical uncertainties are shown as error bars but most are smaller than the symbols.

particles are emitted early from hotter sources and therefore have higher kinetic energies on average. Because of the large binding energy difference between ^3He and ^4He , this effect manifests itself most clearly in the $^3\text{He}/^4\text{He}$ spectra. This dependence of slope on fragment mass makes the isotope ratios sensitive to the range of kinetic energies chosen for the calculation, which in turn influences the isotope-ratio temperature, as will be apparent below.

Figures 2 and 3 also illustrate the expected behavior as a function of excitation energy, i.e., the spectral slopes are flatter at $E^*/A=8$ MeV than at $E^*/A=4$ MeV, implying emission from a hotter source. There is little difference between the forward angle data and that at backward angles. This indicates the minimal impact of fast-cascade/preequilibrium processes in the acceptance gates for the isotope-ratio calculations.

The isotope-ratio temperatures T_{app} corresponding to a given excitation energy per nucleon, were calculated with the double-isotope-ratio thermometer [8]. According to Albergio

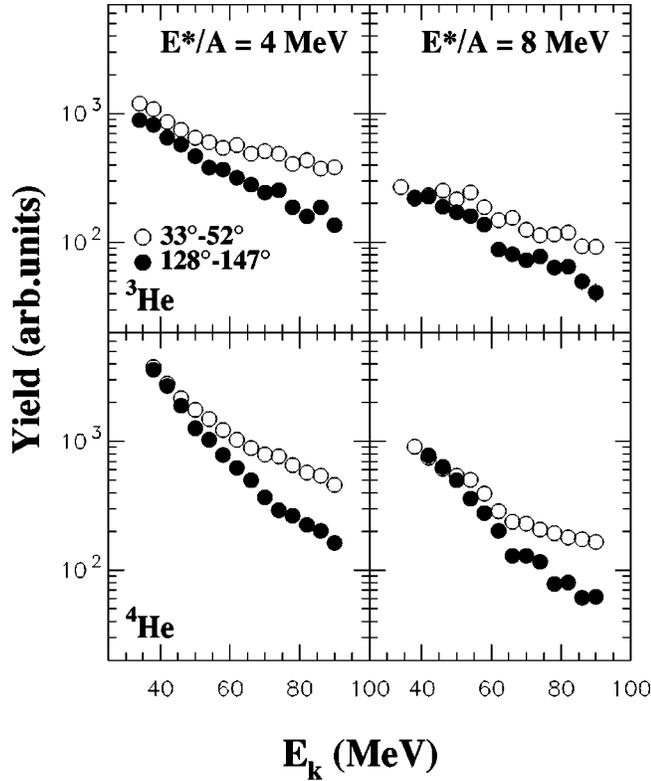


FIG. 3. Forward- and backward-angle kinetic energy spectra for ${}^3\text{He}$ and ${}^4\text{He}$ at $E^*/A=4$ MeV (left panels) and $E^*/A=8$ MeV (right panels). Statistical uncertainties are shown as error bars but most are smaller than the symbols.

et al., the temperature for a system in chemical and thermal equilibrium can be extracted from a double-isotope ratio

$$T_{\text{app}} = \frac{B}{\ln(aR)}, \quad (4)$$

where B is the binding-energy parameter, a is a factor that depends on statistical weights of the ground state nuclear spins, and R is the ground state population ratio at freeze-out. Tsang *et al.* [31] have shown that this method is consistent only when B is greater than about 10 MeV, which in practice means that it is essential to use at least one neutron-deficient isotope. Most current caloric curves use ${}^3\text{He}$ for this isotope.

The p/d , d/t , and ${}^3\text{He}/{}^4\text{He}$ ratios are plotted in Fig. 4 as a function of kinetic energy. The p/d ratios initially decrease up to kinetic energy $E_k \leq 25$ MeV, followed by a slight increase above this kinetic energy. The d/t ratios increase slightly with increasing energy. A much stronger increase in slope is observed for the ${}^3\text{He}/{}^4\text{He}$ ratios, consistent with known importance of preequilibrium ${}^3\text{He}$ emission relative to ${}^4\text{He}$ and the EES predictions [3,30,32,33] discussed above. Thus, the value of the double-isotope-ratios are dominated by the ${}^3\text{He}/{}^4\text{He}$ ratio and it is clear from Eq. (4) that the derived temperature for a given excitation energy will depend on the kinetic energy acceptance of the calculation; the higher the energy cut, the higher the temperature. Unlike the ISiS data of Refs. [12,13], which were obtained with ${}^3\text{He}$

beams, the ratios for the π^- and \bar{p} beams show little sensitivity to angle, permitting use of the full statistics for the caloric curves presented here.

Caloric curves for the 8 GeV/c π^- , $\bar{p} + {}^{197}\text{Au}$ reaction were constructed from two double-isotope-ratios ($R_{dt-\text{He}}$ and $R_{pd-\text{He}}$). The apparent isotopic temperatures are given by

$$T_{\text{app-}dt} = \frac{14.29}{\ln(1.59R_{dt-\text{He}})} \quad (5)$$

and

$$T_{\text{app-}pd} = \frac{18.4}{\ln(5.05R_{pd-\text{He}})}. \quad (6)$$

The kinetic-energy acceptance for R (energy gate to select particles to calculate R) was calculated using ${}^3\text{He}$ and ${}^4\text{He}$ energies between 38–48 MeV. The lower limit was determined by the energy threshold for isotopic resolution of ${}^4\text{He}$ for the ISiS telescopes, set by the energy required to punch through the 500 μm silicon detector. Accounting for the measured Coulomb shifts between $Z=1$ and 2, protons, deuterons, and tritons with energies between 30–40 MeV were taken for the H isotopes in the temperature calculation. The 8 MeV Coulomb shift between $Z=1$ and $Z=2$ is imposed as indicated by the location of the evaporative peaks. The upper limits were based on the balance between statistics and the minimization of preequilibrium effects. As can be seen from Fig. 4 and Eqs. (5) and (6), because of these high energy cuts, T_{app} values will be higher than for studies that employ lower energy H and He ions. The upper panel of Fig. 5 shows the T_{app} heating curves for the 8 GeV/c $\pi^- + {}^{197}\text{Au}$ reaction from the two thermometers, p/d - ${}^3\text{He}/{}^4\text{He}$ and d/t - ${}^3\text{He}/{}^4\text{He}$. There is a difference in the temperature between the two isotopic thermometers at low excitation energy per nucleon that grows in magnitude as the excitation energy increases. Since the ${}^3\text{He}/{}^4\text{He}$ ratio is the same for both thermometers, the deviation for $E^*/A \geq 7$ MeV reflects the fact that in this higher excitation energy region the d/t ratio is smaller than the p/d ratio, as shown in Fig. 4.

Measured yield ratios differ from the primary yield ratios due to sequential decay of the excited fragments. The proton and ${}^4\text{He}$ yields, both of which appear in the numerator of R , are enhanced by evaporation from heavier fragments. Tsang *et al.* [31] have proposed an empirical method to account for these effects by defining a correction factor κ for each isotope ratio by the relationship $R_{\text{app}} = \kappa R_0$, where R_{app} and R_0 are the measured and equilibrium values of the double-isotope-ratio, respectively. From the definition of Eq. (4), the temperature of the fragments at freeze-out (T_0) can be calculated from

$$\frac{1}{T_{\text{app}}} = \frac{1}{T_0} + \frac{\ln \kappa}{B}. \quad (7)$$

Using Tsang's correction factors of $\ln \kappa/B = 0.0497$ and 0.0097 for p/d - ${}^3\text{He}/{}^4\text{He}$ and d/t - ${}^3\text{He}/{}^4\text{He}$, respectively, the caloric curves from the two thermometers are in much better accord at low excitation energy per nucleon ($2 \leq E^*/A$

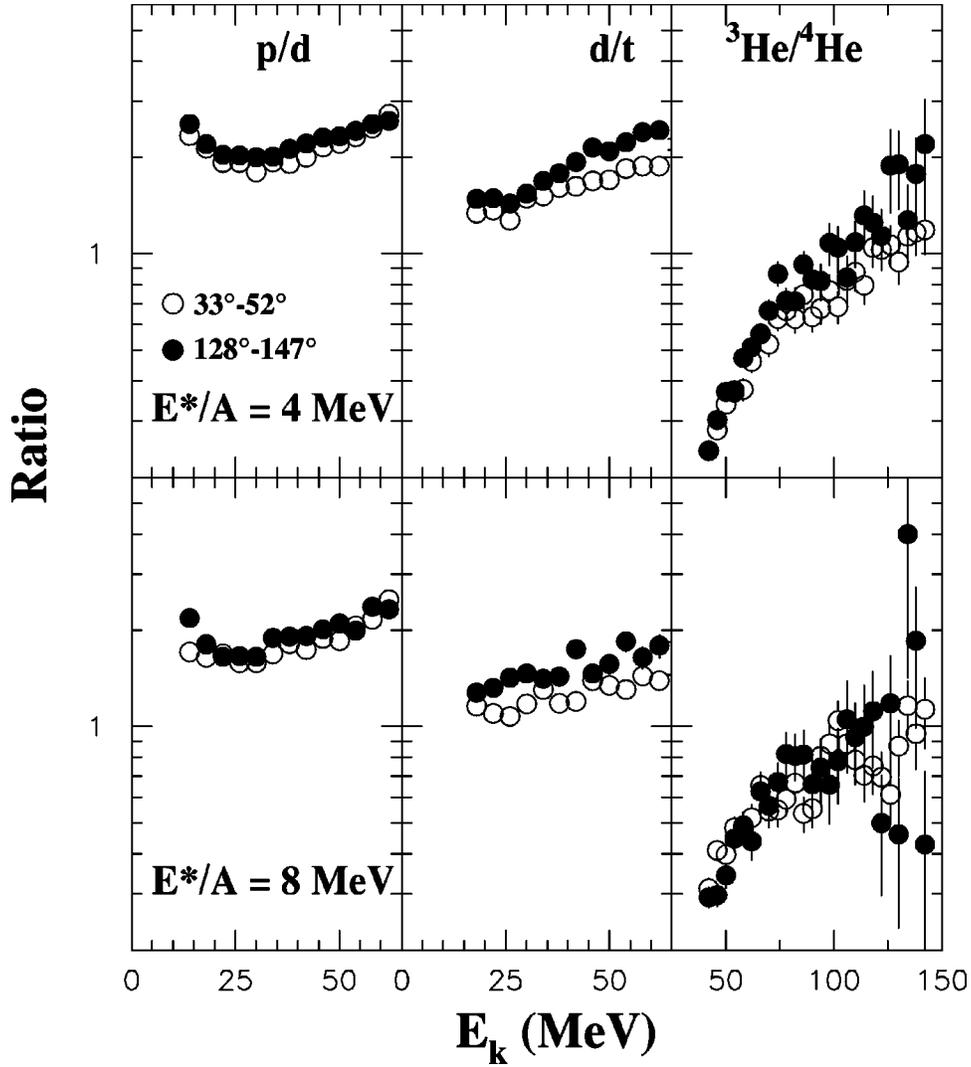


FIG. 4. p/d , d/t , and ${}^3\text{He}/{}^4\text{He}$ isotope yield ratios as a function of kinetic energy in the average source frame for forward-angle (open circles) and backward-angle (solid circles) at $E^*/A=4$ MeV (top panels) and $E^*/A=8$ MeV (bottom panels).

≤ 8), as shown in the bottom panel of Fig. 5. Note that these correction factors are schematic. At higher excitation energies the two thermometers still diverge significantly, a possible consequence of the inability to distinguish between energetic thermal particles and low energy preequilibrium particles at high E^*/A .

The extracted freeze-out temperatures (T_0) increase gradually in the range $1.5 \leq E^*/A \leq 4$ MeV, followed by a flattening of the slope in the region approximately $4 \leq E^*/A \leq 8$ MeV for the d/t - ${}^3\text{He}/{}^4\text{He}$ case and a more distinct plateau in the range $4 \leq E^*/A \leq 11$ MeV for the p/d - ${}^3\text{He}/{}^4\text{He}$ thermometer. At higher excitation energies, the caloric curve from the d/t - ${}^3\text{He}/{}^4\text{He}$ thermometer indicates a more rapid increase in temperature may occur, as in the original ALADIN data [6,7]. The caloric curve from the p/d - ${}^3\text{He}/{}^4\text{He}$ thermometer does not show such a marked increase. Lower statistics for events with $E^*/A \geq 9$ MeV lead to uncertainties for this portion of the curves.

From experiment E900a, the final data set contained 2.5×10^3 antiproton events. The investigation of the heating of

nuclear matter with 8 GeV/c \bar{p} and π^- beams shows significant enhancement of energy deposition in high multiplicity events for \bar{p} compared to π^- [25,26,29], thus producing an excitation energy distribution for \bar{p} that extends to higher values than for π^- . The upper panel of Fig. 6 shows the comparison of the caloric curves derived from the π^- data and \bar{p} data using the d/t - ${}^3\text{He}/{}^4\text{He}$ thermometer. It appears that the caloric curve derived from the \bar{p} data are suggestive of a flatter plateau, although above $E^*/A \sim 3$ MeV, the two curves are statistically the same. However, because of the multiplicity-three trigger level, events with excitation energies less than or equal to 2 MeV have a greater uncertainty due to fluctuations.

A number of observables have been investigated to account for possible differences between the π^- and \bar{p} caloric curves. The residue mass of the thermal source from the experiment is well established, especially at $E^*/A \geq 5$ MeV [29]. The lower panel of Fig. 6 shows the mass of the residue nucleus (Ares) as a function of the excitation energy per

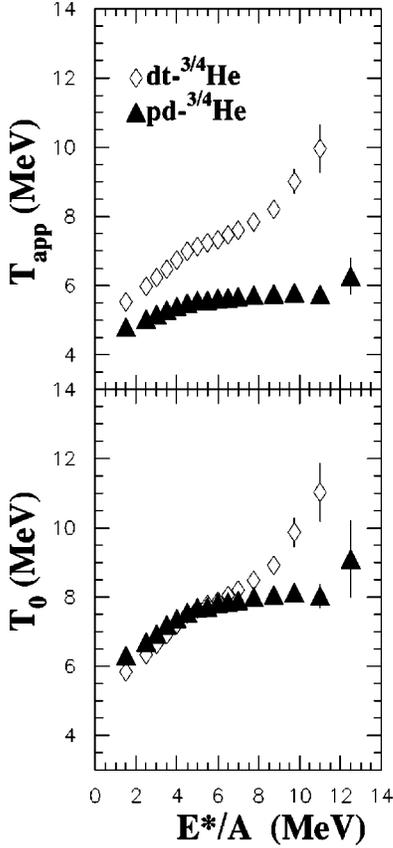


FIG. 5. Caloric curve for 8 GeV/c $\pi^- + ^{197}\text{Au}$ from $p/d-^3\text{He}/^4\text{He}$ and $d/t-^3\text{He}/^4\text{He}$ thermometers using apparent yield to calculate temperature (top panel) and temperature corrected for secondary decay (bottom panel).

nucleon for both π^- and \bar{p} data. The reconstructed masses of the residue nuclei decrease linearly with excitation energy for both π^- and \bar{p} data and agree within 4% for π^- and \bar{p} events. The normalized $Z=1$ and $Z=2$ energy spectra are also very similar for both projectiles, except that spectra from the \bar{p} events have lower statistics.

B. Interpretation

The dependence on the H and He isotope kinetic-energy acceptance (energy gates to calculate double isotope ratio R) for selecting particles for the double-isotope ratio temperature calculation is shown in Fig. 7. These selected gates are independent of the procedure for calculating E^*/A [23]. Here caloric curves from the E900a data are compared using four different kinetic energy gates on the H and He spectra. It can be clearly seen that the higher kinetic energies gates yield higher temperatures from the double-isotope ratio. This behavior is a direct consequence of the increase in the $^3\text{He}/^4\text{He}$ ratio as a function of He energy (Fig. 4), which is the dominant ratio in determining the temperature. Similar behavior for the $^3\text{He}/^4\text{He}$ ratio is observed in lower energy hadron-induced reactions [34–36]. The caloric curves from the three lowest energy gates (the open circles, the solid triangles, and the open triangles) have similar systematic

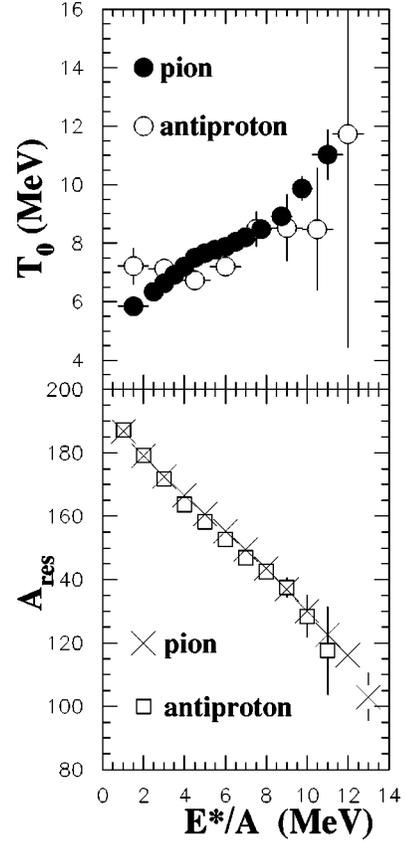


FIG. 6. Top panel shows the comparison of caloric curves using $d/t-^3\text{He}/^4\text{He}$ thermometer corrected for secondary decay from 8 GeV/c $\pi^- + ^{197}\text{Au}$ (solid circles) and 8 GeV/c $\bar{p} + ^{197}\text{Au}$ (open circles) reactions. Bottom panel compares the residue mass of pion data (crosses) with antiproton data (open squares). Statistical uncertainties are shown as error bars but most are smaller than the symbols.

trends. All three caloric curves show an increase in temperature in the interval $1 \geq E^*/A \geq 5$ MeV followed by a slope change at higher excitation energies. The slope change, which may indicate a phase transition, disappears for the highest energy gate.

The sequence of caloric curves in Fig. 7 can be interpreted as evidence for “cooling” of the hot residues as they evolve from the fast cascade stage of the collision toward equilibrium [12,37]. In this preequilibrium/coalescence regime, early emission favors the production of more energetic particles and simpler clusters (e.g., ^3He relative to ^4He as in Fig. 4). The net effect is to provide time-dependent snapshots of the “cooling curve” for the system. Alternatively, this cooling behavior can be described within the framework of the statistical EES model [3], since each particle spectrum consists of a convolution of Maxwellian distributions and the hottest sources are the most significant contributions to the high energy tails of the spectra.

From the systematic trends shown in Fig. 7 it is estimated that the isotope ratio temperatures would be lower if extrapolated to the evaporative H and He peaks (Fig. 1). For each excitation energy bin, a linear fit was made to temperature versus kinetic energy. This was then extrapolated to lower

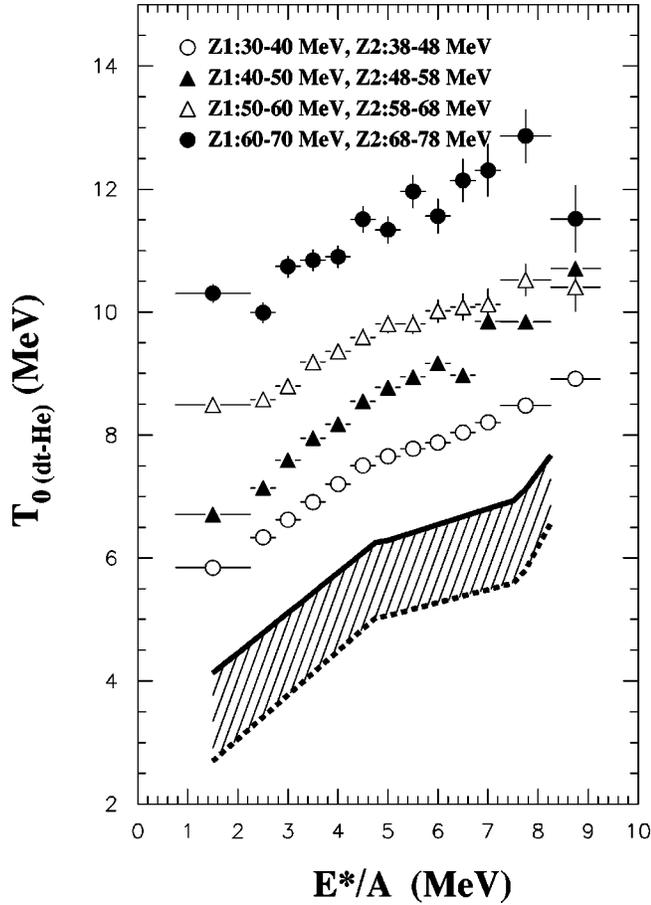


FIG. 7. The caloric curve for $8 \text{ GeV}/c \pi^- + {}^{197}\text{Au}$ from the $d/t\text{-}^3\text{He}/^4\text{He}$ thermometer corrected for secondary decay using four different kinetic energy acceptances as given in the graph in MeV. The shaded area represent the region of caloric curves that extrapolated to the evaporative H and He peaks.

kinetic energy. The solid and dashed lines in Fig. 7 show such extrapolations for two lower energy acceptance gates [$Z=1$: 20–30 MeV, $Z=2$: 28–38 MeV (solid line) and $Z=1$: 10–20 MeV, $Z=2$: 18–28 MeV (dashed line)]. These provide a reasonable upper and lower bound for the caloric curve. This comparison illustrates the sensitivity of temperatures from the double-isotope-ratio thermometer to the energy acceptance of each isotope pair, and emphasizes that the T_0 values determined in this work would be lower if isotopic resolution in ISiS could be obtained at lower kinetic energies.

Figure 8 compares the caloric curves from various experiments with a ${}^{197}\text{Au}$ target, which use the double-isotope ratio based on ${}^3\text{He}$ to extract the temperature and the calorimetry method to calculate the excitation energy. Here, the caloric curves from E900a experiments using $p/d\text{-}^3\text{He}/^4\text{He}$, $d/t\text{-}^3\text{He}/^4\text{He}$ thermometers and the extrapolated caloric curves (same as Fig. 7) are compared with caloric curves from other experiments. In addition to the present data, the plots include data from the ISiS E228 experiment ($4.8 \text{ GeV } {}^3\text{He} + {}^{197}\text{Au}$) [12,13], the EOS Collaboration ($1 \text{ GeV } {}^{197}\text{Au} + \text{C}$) [9–11], and the ALADIN group ($600 \text{ MeV/nucleon Au} + \text{Au}$) [6,7]. The curves all fit into a broad

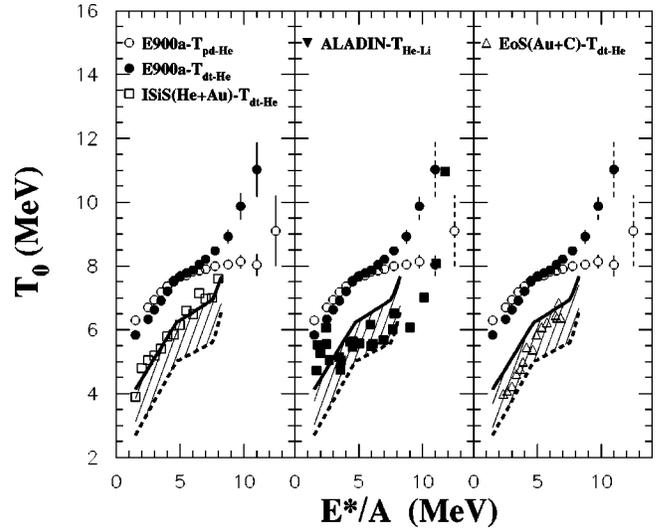


FIG. 8. Caloric curves from the E900a experiment ($8 \text{ GeV}/c \pi^- + {}^{197}\text{Au}$, same data as in bottom panel of Fig. 5), ISiS E228 experiment (Refs. [12,13]), ALADIN data (Refs. [6,7]), and EOS data (Refs. [9–11]).

band between the two extrapolated caloric curves in the temperature-excitation-energy plane. When the caloric curves are extrapolated to lower energy acceptance gate, good agreement is observed with the ALADIN data the EOS data.

Another difference between the EOS and ISiS caloric curves is the methodology used for the excitation energy calculation [23]. To investigate the effect of the selection criteria for thermal particles in the E^*/A calculation, the excitation energies from E900a were recalculated using the same energy selection for thermal particles as the EOS experiment [9–11]. The EOS kinetic energy acceptance for thermal particles is defined as all particles with energies less than $E_k/A \leq 30 \text{ MeV}$, which yields higher excitation energies than the ISiS prescription for E^* by about 20% [23]. Figure 9 shows the caloric curves from E900a using this EOS excitation energy definition compared with the E900a caloric curve in the lower panel of Fig. 5 and the Au+C EOS caloric curve. The resulting ISiS caloric curve is shifted to higher excitation energy, thus producing a flatter and lower caloric curve than that adopted here. The lower T_0 values for the EOS relative to the E^* -adjusted ISiS data result in part from the lower kinetic energy acceptance for isotope identification in the EOS experiment and it agrees well with the extrapolated curve shown in Fig. 7.

Recent work by Chomaz *et al.* [38] investigated the caloric curves and energy fluctuations in a microcanonical liquid-gas phase transition and found that caloric curves may differ depending on the path followed in the thermodynamical variable plane. The size of the systems, which depends on collision dynamics, may also affect the curve. Natowitz [39] has recently investigated a variety of caloric curves and concluded the measurements are self-consistent when the size of the system is taken into account. Thus many reasons may be responsible for the difference in the caloric curves. Each experiment has different formation mechanisms, fragment ki-

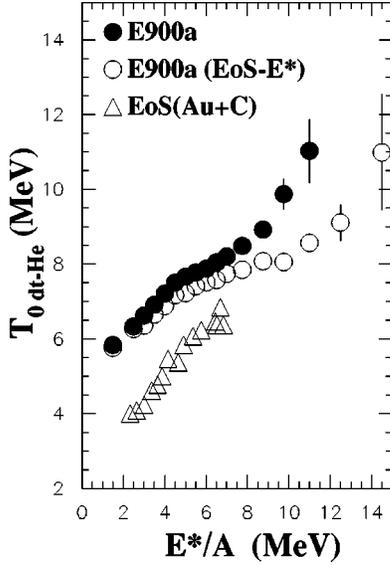


FIG. 9. Caloric curve from the $8 \text{ GeV}/c \pi^- + {}^{197}\text{Au}$ reaction using the $d/t\text{-}^3\text{He}/^4\text{He}$ thermometer corrected for secondary decay (solid circles); using the EoS excitation energy calculation (open circles); and from EoS Collaboration experiment (open triangles).

netic energy acceptance, and particle identification limitations due to instrumentation and analysis procedures. As illustrated for the ISiS data in Fig. 7, since the kinetic energy cuts for the calculation of R in Eq. (4) are higher than for the other experiments, the ^3He contribution, and therefore T_{app} , are higher.

IV. DISCUSSION

In addition to the caloric curve behavior that has been explored in this and other studies, several other signals expected of a liquid-gas phase transition have been reported. Among these are very short disintegration times [40], negative heat capacity [41], and percolation [42]/Fisher scaling [43] analyses of the IMF multiplicity distributions that indicate critical behavior. If we take the total of these results at face value, then the important question remains: is the phase transition in hot nuclei first-order or continuous?

In Fig. 10, we show the relative emission time scale for this reaction [40] overlaid on the $\pi^- + {}^{197}\text{Au}$ caloric curve in the lower frame, along with the charge-distribution power-law parameter (τ) [29] in the upper frame, where $\sigma_Z \propto Z^{-7}$. These results should be viewed in the context of Fisher scaling [43] and percolation analyses [42], both of which show evidence for critical behavior near $E^*/A \sim 4\text{--}5 \text{ MeV}$. In the bottom panel of Fig. 10, a distinct change in the slope of the caloric curve is observed between $E^*/A \sim 4\text{--}8 \text{ MeV}$, suggestive of but not fully consistent with the behavior expected for a first order phase transition. The beginning of this slope change region occurs at just the point where the emission time scale achieves a minimum, the probability for large cluster size reaches a maximum, and evidence for critical behavior is observed. An alternative possible interpretation of Fig. 10 is that once the system reaches the critical energy, its density continues to decrease

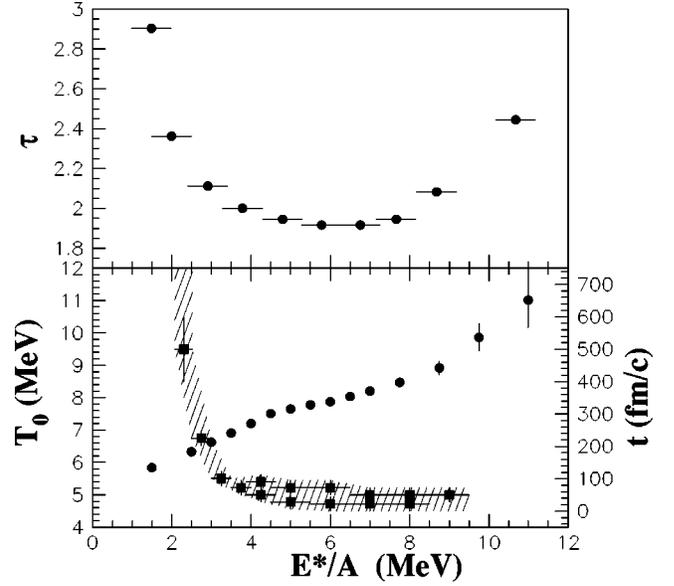


FIG. 10. The top panel shows the charge-distribution power-law parameter as a function of excitation energy per nucleon from Ref. [29]. The bottom panel shows the caloric curve using the $d/t\text{-}^3\text{He}/^4\text{He}$ thermometer, corrected for secondary decay from the $8 \text{ GeV}/c \pi^- + {}^{197}\text{Au}$ reaction overlaid on the relative emission time scale for the reaction from Ref. [40].

with E^*/A , keeping T constant and implying that the system remains in the vicinity of the critical point over a significant range of excitation energy.

It is also instructive to compare these caloric curve data with nuclear models that assume a phase transition. In Fig. 11, the caloric curves from E900a are compared to predictions from the statistical multifragmentation model (SMM) [44], using the $d/t\text{-}^3\text{He}/^4\text{He}$ thermometer. The dashed line and the solid line are the extrapolated caloric curves (same as in Fig. 7). The experimentally determined residue charges and masses corresponding to the excitation energy were fed as input to the SMM code and gates were imposed on the calculated spectra to coincide with the experimental acceptance. Figure 11 shows the experimentally derived apparent temperatures compared with the SMM prediction. The solid square symbols represent the prediction when running the “cold” version of the code. In the “cold” fragmentation scenario, the fragments remain cold or weakly excited in the course of the fragmentation process [1]. In this case, the breakup of the system leaves the emitted fragment in a low state of internal excitation, thus suppressing secondary decay. Here the temperatures were calculated with the $d/t\text{-}^3\text{He}/^4\text{He}$ ratio using the same energy acceptance as in the experiment, i.e., H at 30–40 MeV and He at 38–48 MeV. For the “cold” mode, the correction factor for secondary decay is not applied. To see the effect of the energy acceptance on the SMM predictions, the caloric curve is also derived from the double isotope ratio temperatures by taking all particles that have kinetic energies greater than 5 MeV, represented by open squares. This caloric curve shows more plateaulike behavior for $3 \leq E^*/A \leq 6 \text{ MeV}$ and has lower temperatures at higher E^*/A than the caloric curve that im-

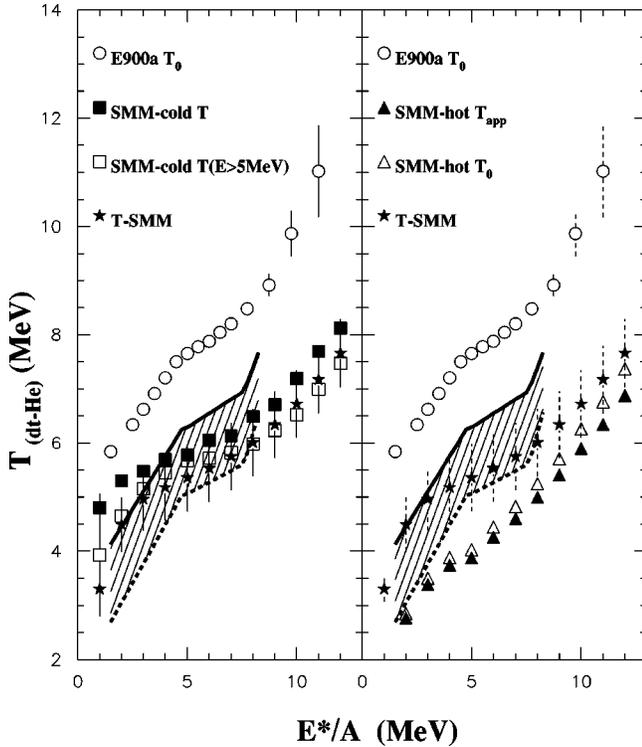


FIG. 11. The caloric curve using the $d/t\text{-}^3\text{He}/^4\text{He}$ thermometer, corrected for secondary decay from the 8 GeV/c $\pi^- + {}^{197}\text{Au}$ reaction, (open circles); “cold” version of SMM prediction using the $d/t\text{-}^3\text{He}/^4\text{He}$ thermometer and imposed the same energy acceptance as in the experimental data (solid squares); “cold” SMM version without the imposed energy acceptance (open squares); “hot” SMM version without the correction factor for secondary decay (solid triangles); and with correction for secondary decay (open triangles); and the SMM prediction from the microcanonical temperature (stars).

poses the same energy acceptance as the experimental data.

For comparison, the SMM code was also run in the “hot” mode, for which some of the available excitation energy resides in the fragments. Then these prefragments can undergo secondary decay [1]. This has the effect of enhancing proton and ^4He yields and lowering T_{app} . The solid triangles are the caloric curve prediction from the “hot” version of the code. The open triangle represents the caloric curve of the “hot” mode of SMM in which the correction factor for secondary decay was applied in the double-isotope-ratio calculation. It can be clearly seen that the magnitude and shape of the “cold” version of SMM are in better agreement with the experimental result than the “hot” version, especially when one accounts for the “cooling effect” discussed earlier, the data are more consistent with the cold version of SMM. The microcanonical temperatures in the SMM calculation are shown by star symbols. The caloric curve from microcanonical temperatures is in better agreement with “cold” mode than “hot” mode. This implies that isotopic temperatures can be strongly affected by the secondary decays of hot primary fragments.

However, the SMM does not take into account the light charged particles that are emitted during the thermal expan-

sion of the nuclear system prior to fragment formation [45]. Therefore, the final model analysis should involve a comparison of the energy spectra of the emitted particles.

V. SUMMARY

In summary, hadron-induced reactions provide transparent systems for investigating the thermal properties of highly excited nuclei, with minimal contributions from dynamical effects such as compression and rotation. This paper presents caloric curves from the 8 GeV/c π^- and $\bar{p} + {}^{197}\text{Au}$ reactions. The caloric curves have been derived from reconstructed excitation energies and two double-isotope-ratio temperatures $d/t\text{-}^3\text{He}/^4\text{He}$ and $p/d\text{-}^3\text{He}/^4\text{He}$. Contributions to the temperatures from secondary decay have been accounted for with empirical correction factors [31]. The caloric curves from the two thermometers ($p/d\text{-}^3\text{He}/^4\text{He}$ and $d/t\text{-}^3\text{He}/^4\text{He}$) overlap below $E^*/A \sim 8$ MeV when the correction factors are applied, but diverge significantly at higher excitation energy. The failure of the two temperatures to converge at the higher excitation energies indicates that more work is needed to understand the effects of both the cascade process and secondary decay for the higher excitation energies.

The freeze-out temperatures (T_0) increase gradually in the temperature range $1.5 \leq E^*/A \leq 4$ MeV, followed by a slope change at approximately $4 \leq E^*/A \leq 8$ MeV, where the emission time scale becomes very short, IMF multiplicities increase rapidly and extra expansion energy is observed. An increase in temperature at higher excitation energy, where statistics are low and fluctuations are high, is also observed, in line with a phase transition interpretation. The double-isotope-ratio temperature depends strongly on the kinetic energies of the isotopes that are included in the isotope ratio calculation. As higher energy particles are included, higher temperatures result from the calculation, in essence, providing a time-dependent cooling curve for the hot residues as they evolve toward equilibrium [3,37]. Although the overall temperatures depend on the choice of energy acceptance, the basic character of the caloric curve remains unchanged.

In addition to the kinetic energy acceptances imposed on the spectra, the definition of thermal particles in the E^* calculation can result in different caloric curves. Since each measurement imposes a separate set of acceptances, the caloric curves from different experiments do not totally overlap with one another. However, they all fit in to a broad band on the temperature-excitation-energy plane. When comparing with other data, one must take into account any differences in energy acceptance of both the particles included in the calculation of the temperature and those included in the excitation energy calculation. The effect of these experimental parameters is significant, as has been demonstrated by analysis of the current data.

Overall E900 has produced high statistics data with good energy resolution that represent the highest excitation energy data for large systems ($A \geq 100$) [39]. The data are minimally affected by compression, angular momentum or flow and reach high excitation energies with moderate changes in system size. While the initial rise and plateau have been seen in many systems, the second rise above $E^*/A \sim 9$ MeV may

be of significant interest. This rise is consistent with a phase transition over a narrow range in E^*/A and is complemented by the Fisher droplet analysis [42] and percolation analysis [41], which also find a phase transition and evidence for critical behavior in this region.

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